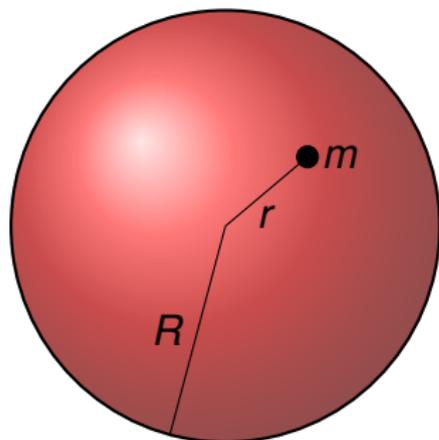
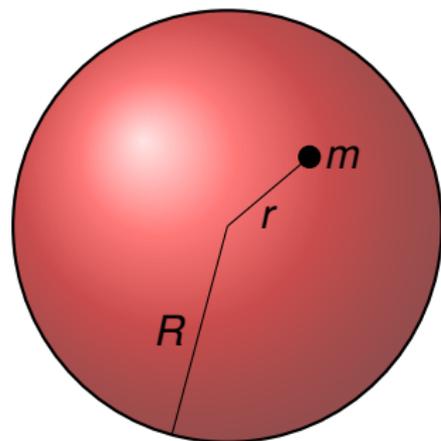


## Example: A mass Inside a Spherical Mass



The shown sphere has a radius  $R$  and a mass  $M$  uniformly distributed over its surface. Another mass  $m$  is placed at a distance  $r < R$  from the center. What will be the gravitational force that the object will feel?

# Example: A mass Inside a Spherical Mass



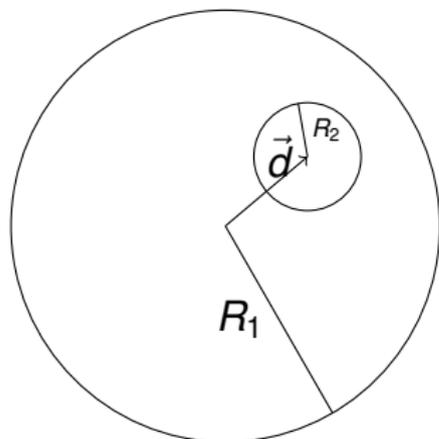
- Divide the sphere into an inner sphere with radius  $r$  and outer shell.
- Outer shell will not exert any force.
- Inner shell will have a mass:

$$M(r) = \frac{M}{\frac{4}{3}\pi R^3} \frac{4}{3}\pi r^3 = M \left(\frac{r}{R}\right)^3$$

- The force exerted by the inner shell is

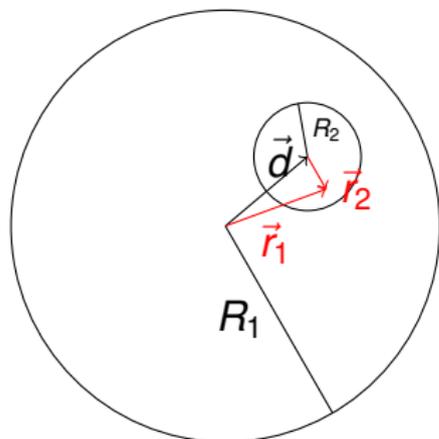
$$\vec{F} = -G_N M \left(\frac{r}{R}\right)^3 \frac{m}{r^2} \hat{r} = -G_N \frac{Mm}{R^3} \vec{r}$$

# Example: A sphere with another sphere carved out



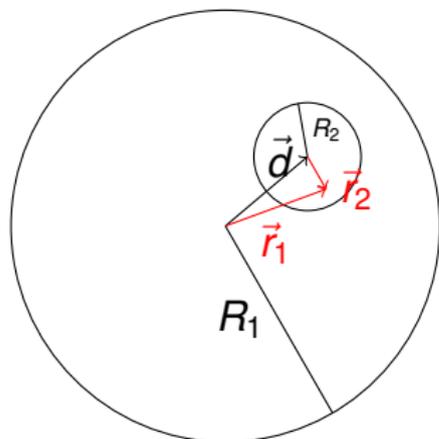
A sphere of radius  $R_2$  is carved out of another sphere of radius  $R_1$ . The position of the center of the carved sphere is denoted by  $\vec{d}$ . The mass density of the system is  $\rho$ . If a mass  $m$  is placed inside the cavity, what will be the force that this mass  $m$  will feel?

# Example: A sphere with another sphere carved out



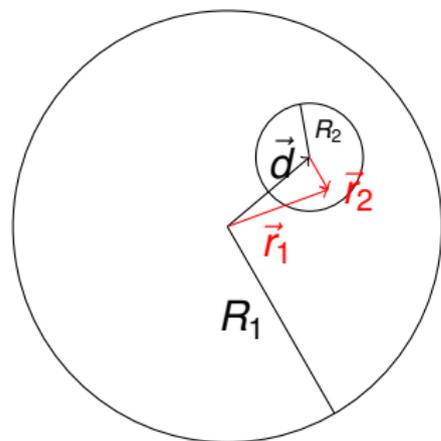
- Let  $\vec{r}_1$  ( $\vec{r}_2$ ) be the position of the mass  $m$  relative to the center of the large (small) sphere.

# Example: A sphere with another sphere carved out



- Let  $\vec{r}_1$  ( $\vec{r}_2$ ) be the position of the mass  $m$  relative to the center of the large (small) sphere.
- $\vec{F}_{full\ sphere} = \vec{F}_T + \vec{F}_{carved\ out\ mass} \implies \vec{F}_T = \vec{F}_{full\ sphere} - \vec{F}_{carved\ out\ mass}$
- The cavity can be modeled as a mass with mass density  $-\rho$ .

# Example: A sphere with another sphere carved out



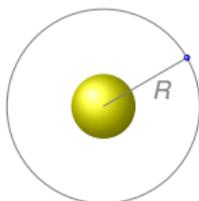
- $\vec{F}_{full\ sphere} = \vec{F}_T + \vec{F}_{carved\ out\ mass} \implies \vec{F}_T = \vec{F}_{full\ sphere} - \vec{F}_{carved\ out\ mass}$
- The cavity can be modeled as a mass with mass density  $-\rho$ .
- Large sphere:
 
$$\vec{F}_L = -G_N \frac{\rho \frac{4}{3}\pi r_1^3}{r_1^2} \hat{r}_1 = -\frac{4\pi}{3} G_N \rho \vec{r}_1$$
- Small sphere:  $\vec{F}_s = -\frac{4\pi}{3} G_N (-\rho) \vec{r}_2$

$$\vec{F}_T = \vec{F}_L + \vec{F}_s = -\frac{4\pi}{3} G_N \rho (\vec{r}_1 - \vec{r}_2) = \frac{4\pi}{3} G_N \rho \vec{d}$$

Gravitational attraction is uniform inside the cavity.

CHALLENGE: Can you prove this without using vectors? (not recommended)

## Example: Circular Orbits



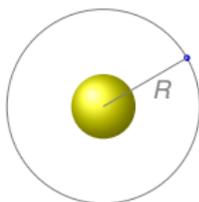
- Let  $M_E$  ( $M_S$ ) be the mass of Earth (satellite)
- What is the speed of the satellite?
- $m \frac{v^2}{R} = G_N \frac{Mm}{R^2} \implies v = \sqrt{\frac{G_N M}{R}}$
- The closer the satellite is to the Earth, the faster it should be.

- The period of the satellite is

$$T = \frac{2\pi R}{v} = \frac{2\pi}{\sqrt{G_N M}} R^{\frac{3}{2}} \implies \frac{T^2}{R^3} = \frac{2\pi}{\sqrt{G_N M}}$$

- Measuring the ratio  $T^2/R^3$ , it is possible to determine the mass of the sun.

## Example: Circular Orbits



- Let  $M_E$  ( $M_S$ ) be the mass of Earth (satellite)
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- Geocentric orbits are those for which the relative position of the satellite is fixed with respect to the surface of the planet.